Probing (Beyond) the Standard Model physics On heaven & earth

Future vision

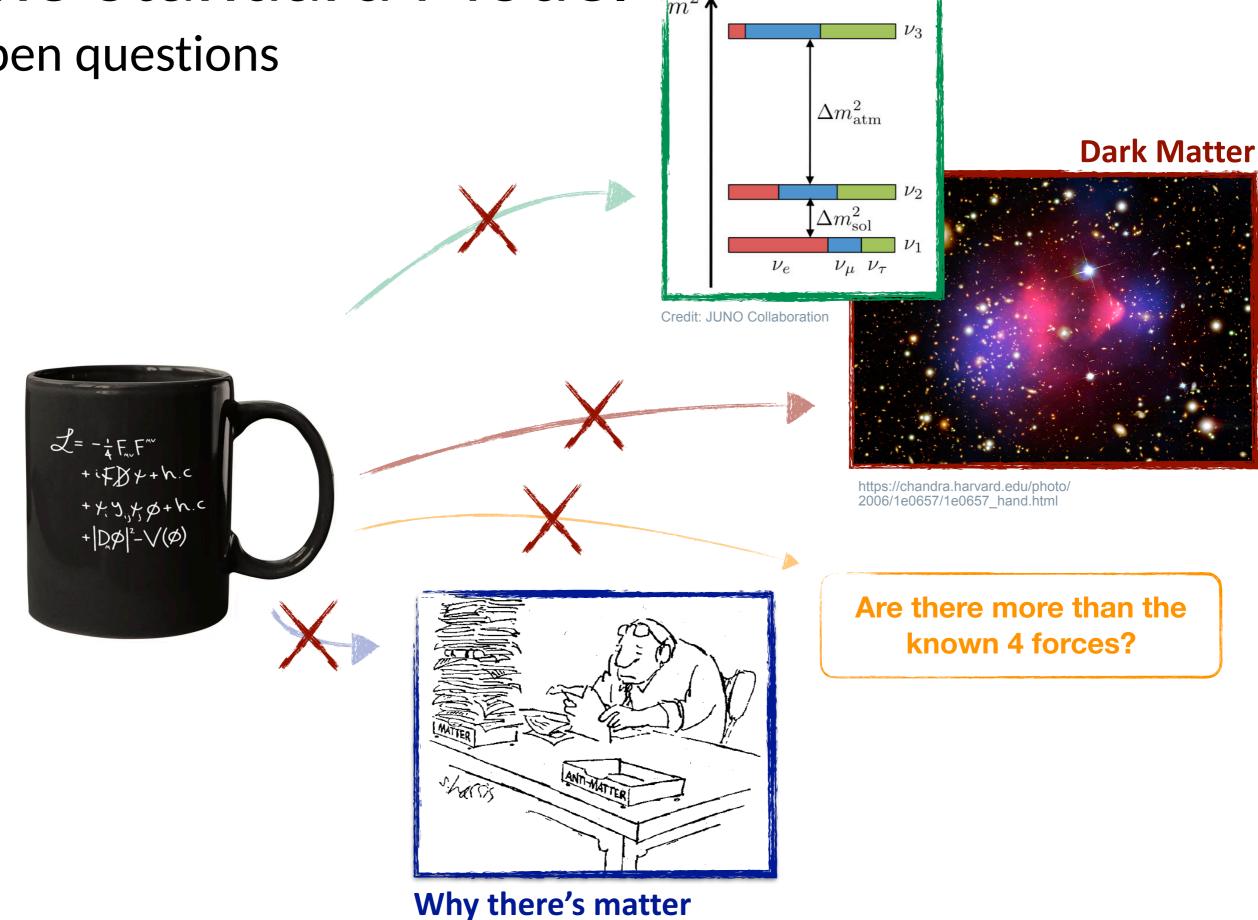
Wouter Dekens





The Standard Model

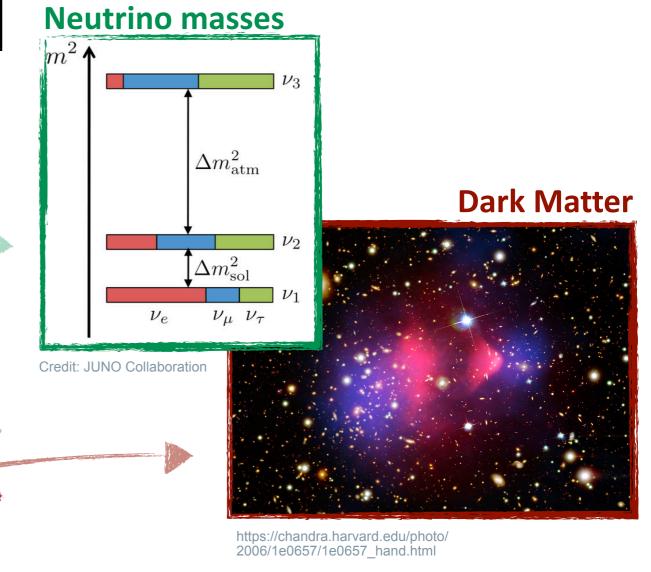
Open questions



Neutrino masses

The Standard Model

Open questions



Answers require new particles and interactions

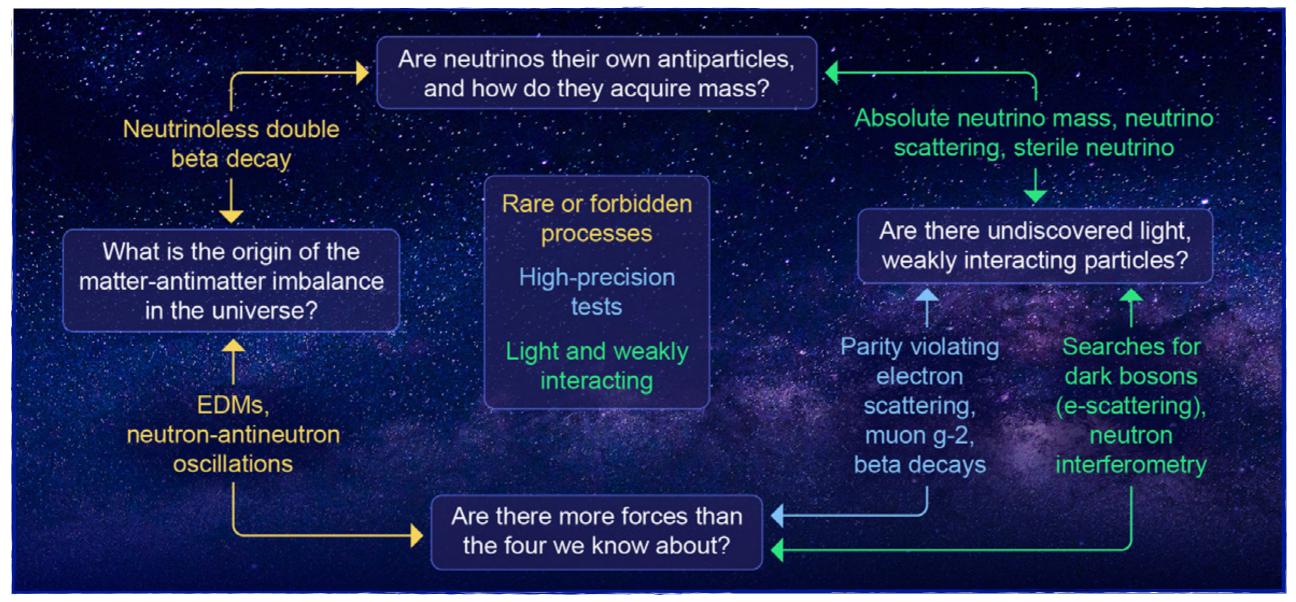




Why there's matter

Are there more than the known 4 forces?

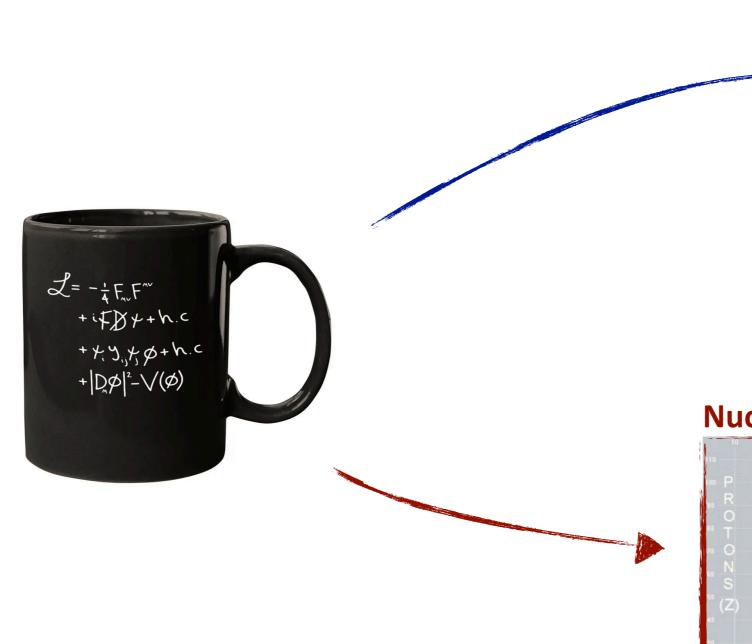
How to find new physics

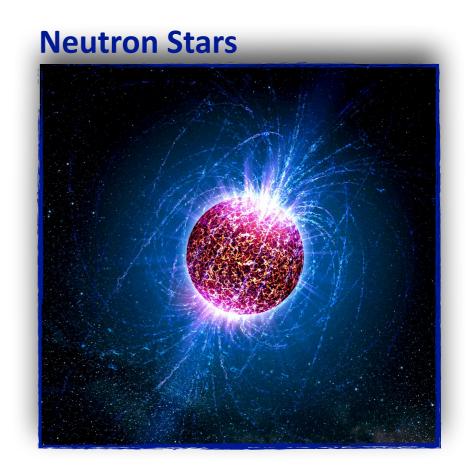


2023 Long Range Plan for Nuclear Science

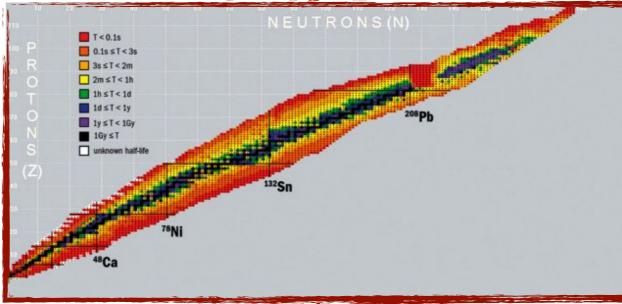
The Standard Model

Hard questions





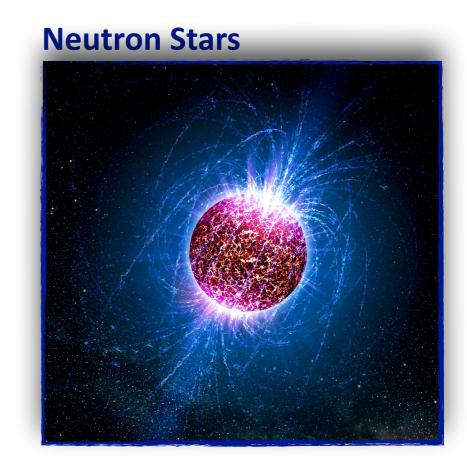
Nuclei



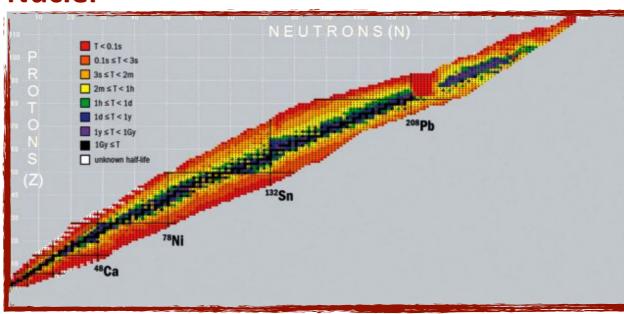
The Standard Model

Hard questions

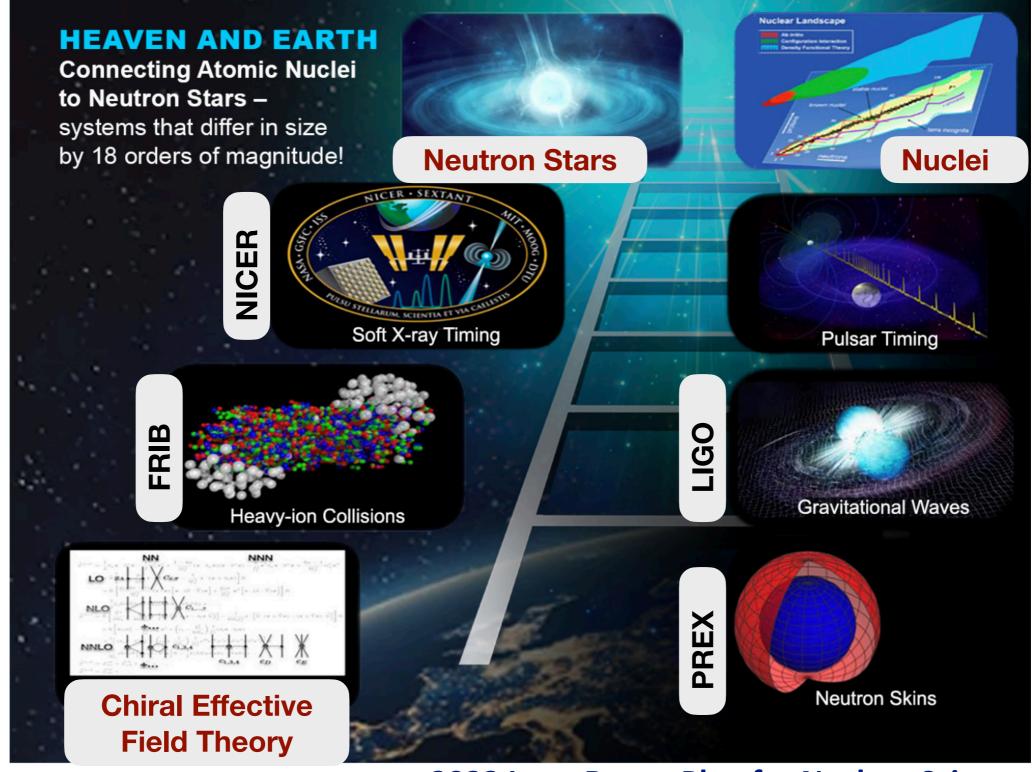
Requires accurate description of Nuclear Forces



Nuclei



Connecting nuclei to the cosmos

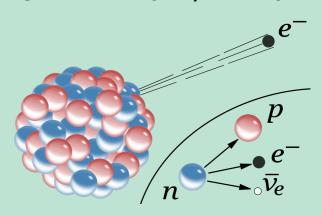


2023 Long Range Plan for Nuclear Science



Precision frontier

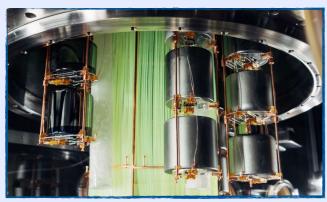
ullet Cabibbo Angle Anomaly & eta decays



Effectively probing (Beyond) the SM physics

Fundamental Symmetries

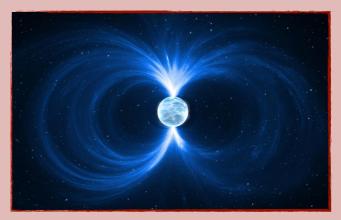
• Lepton number & Sterile Neutrinos

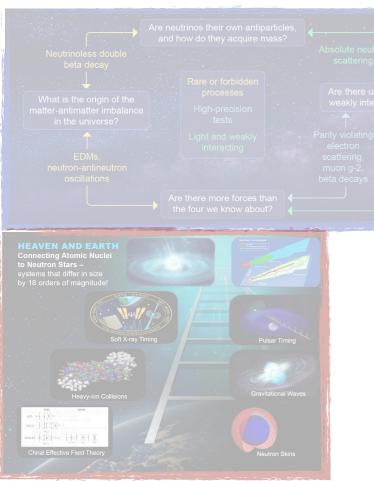


 $0\nu\beta\beta$ LEGEND (ORNL lead US lab)

Connecting nuclei to the cosmos

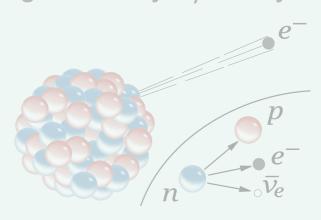
• Dense matter & Neutron stars







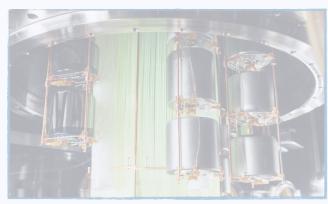
ullet Cabibbo Angle Anomaly & eta decays



Effectively probing (Beyond) the SM physics

Fundamental Symmetries

• Lepton number & Sterile Neutrinos



 $0\nu\beta\beta$ LEGEND (ORNL lead US lab)

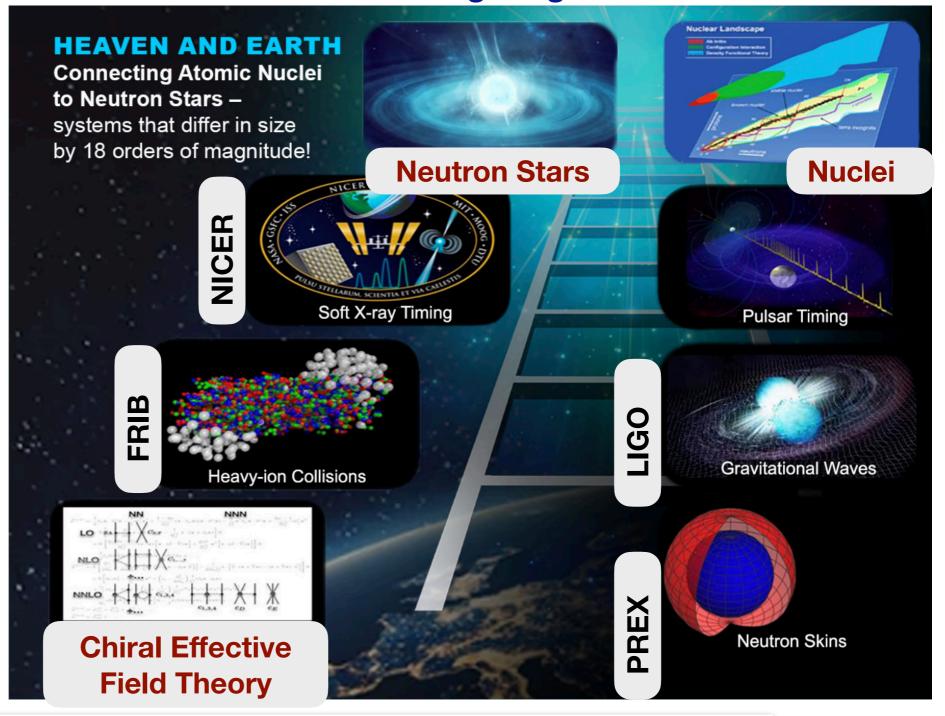
Connecting nuclei to the cosmos

Dense matter & Neutron stars



Nuclear forces

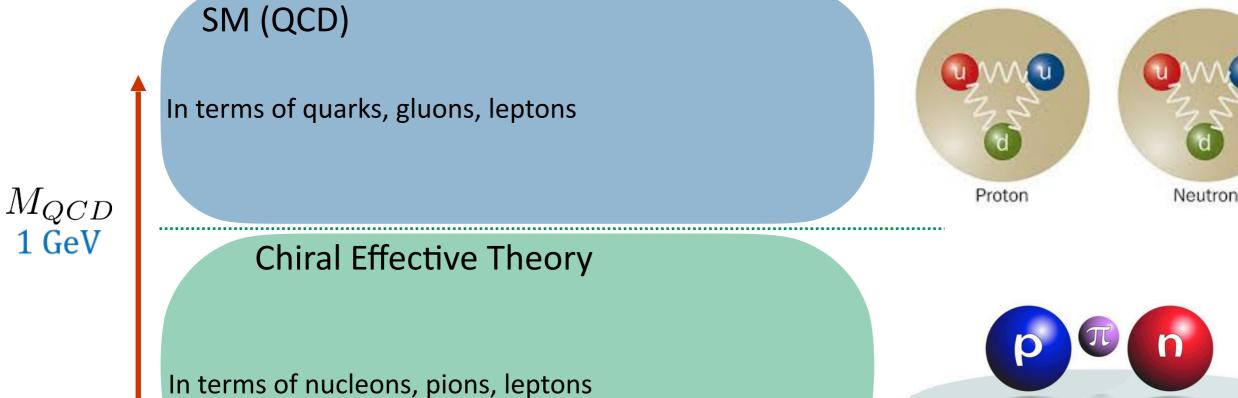
2023 Long Range Plan for Nuclear Science



- Upcoming laboratory & astrophysics experiments will probe nuclear physics of
 - Neutron Stars
 - Nuclei
- Connection requires understanding of nuclear forces

From quarks to nucleons

Modern view of nucleon forces



• Possible hadronic interactions determined by (chiral) symmetries

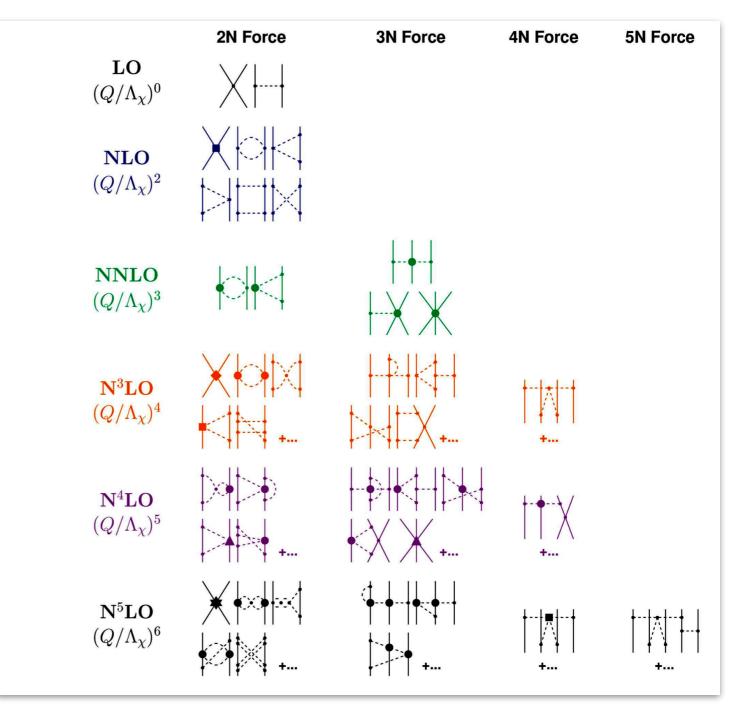
baryons, mesons

- Interaction strengths determined by (unknown) coupling constants
- \bullet Expansion in $Q/\Lambda_{\chi} \sim m_{\pi}/\Lambda_{\chi}$, where $\Lambda_{\chi} \sim 1~{\rm GeV}$

Chiral Effective Theory Modern view of nucleon forces

- Chiral EFT is an expansion in
 - $Q/\Lambda_\chi \sim m_\pi/\Lambda_\chi$, where $\Lambda_\chi \sim 1~{\rm GeV}$
- Leads to a hierarchy of nuclear forces
- Relies on "power counting"
 - Determines the importance of vertices
- Weinberg's counting is widely used
 - Known to have inconsistencies

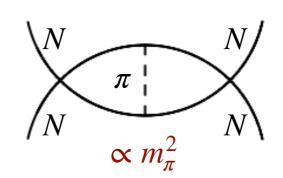
Requires re-ordering three-nucleon forces

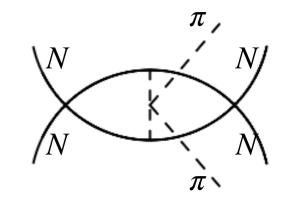


Hierarchy of nuclear forces up to N5 LO in ChiPT. Solid lines represent nucleons and dashed lines pions. Entem, Machleidt, Y. Nosyk, (arXiv:1703.05454)

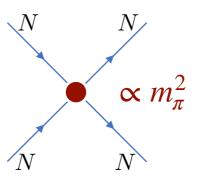
New class of three-nucleon forces Failure of Weinberg's counting

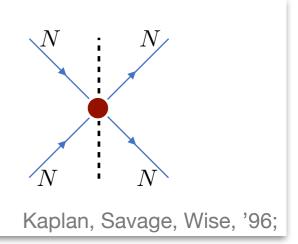
- Certain diagrams lead to divergences
 - Generate short-distance interactions



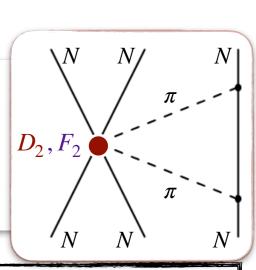


- Consistency requires leading-order interactions
 - N2LO in Weinberg's counting



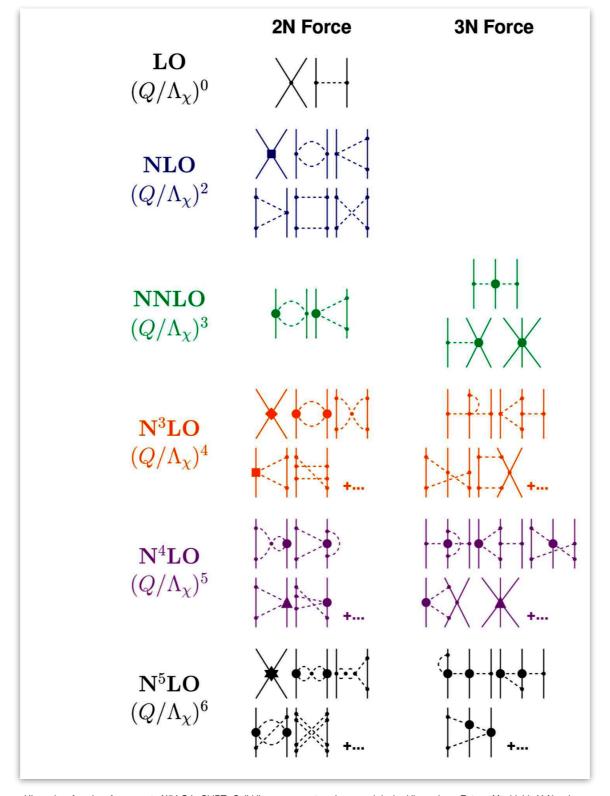


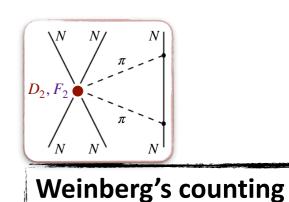
- These interactions generate a new class of 3N forces
- Cannot be absorbed into LO contact terms like in 2-nucleon sector



WD, M. Dawid, V. Cirigliano, S. Reddy

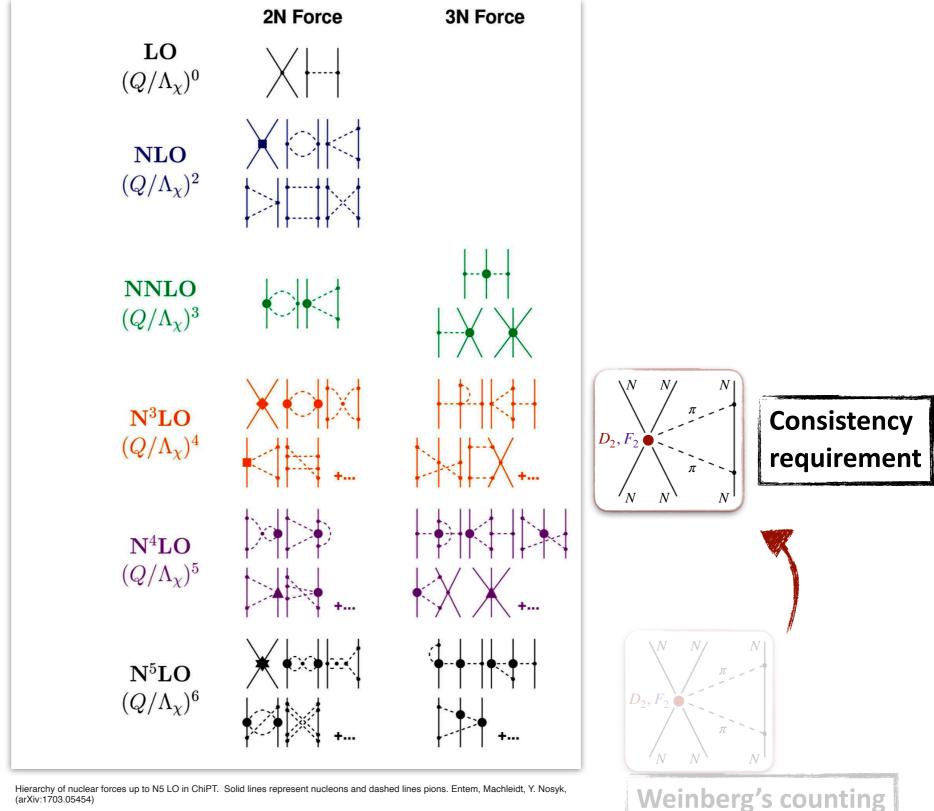
New class of three-nucleon forces Relearning how to count





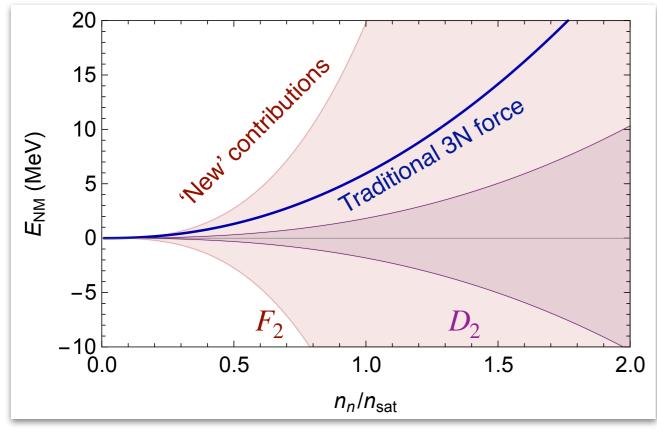
Hierarchy of nuclear forces up to N5 LO in ChiPT. Solid lines represent nucleons and dashed lines pions. Entem, Machleidt, Y. Nosyk, (arXiv:1703.05454)

New class of three-nucleon forces Relearning how to count

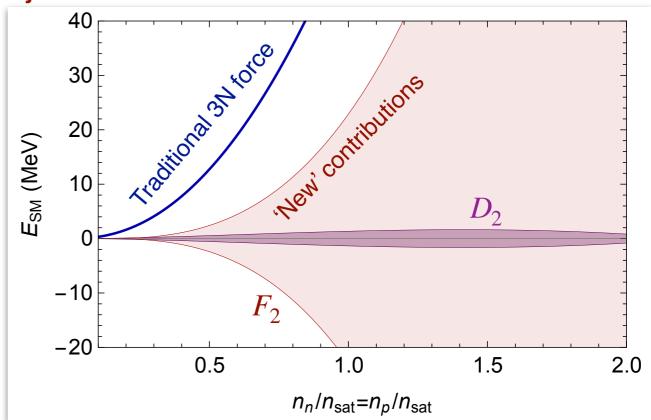


New class of three-nucleon forces Effects in dense matter

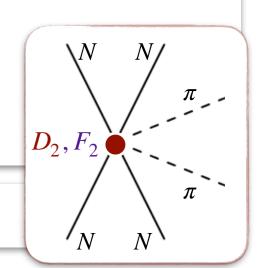
Neutron Matter



Symmetric Matter



- Significant contribution in Neutron & symmetric matter
 - Larger than currently quoted uncertainties
- Affects our understanding of nuclei and Neutron Stars
 - Equation of State in Neutron stars & mass-radius relation
- Depends on poorly known 2pion-4nucleon couplings



From Nuclei to the Cosmos Future projects

Connecting nucleons to neutron stars

• Fit to light nuclei Output Output

Impact on dense matter

- Nuclei
- Neutron Star Equation of State



Spin-off: Effect on axions

Affected by the same hadronic couplings in dense matter

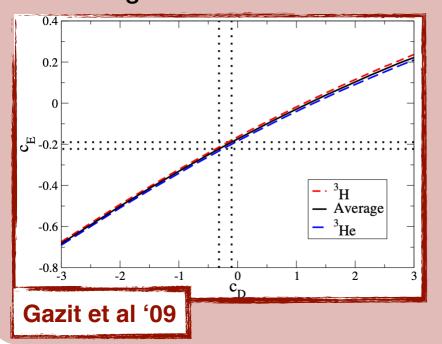


Connecting nucleons to neutron stars

Overlap with Nuclear Theory Group

Determining the hadronic couplings

• Fit to light nuclei



Impact on dense matter

- Nuclei
- Neutron Star Equation of State

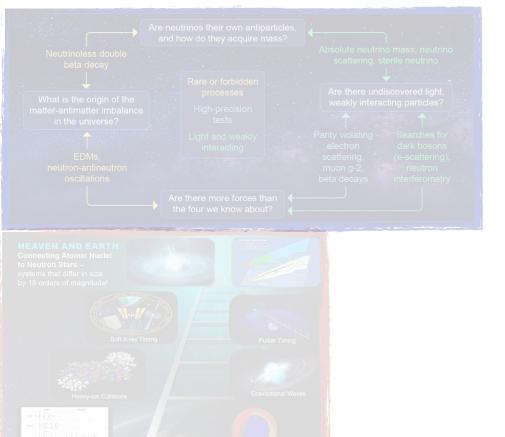


Spin-off: Effect on axions

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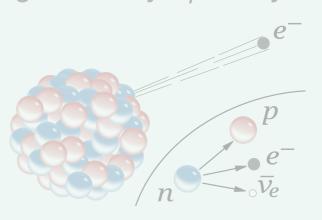


Astrophysics Group



Precision frontier

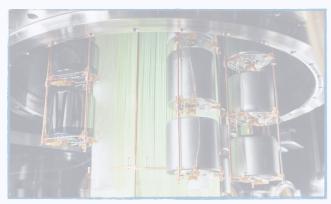
ullet Cabibbo Angle Anomaly & eta decays



Effectively probing (Beyond) the SM physics

Fundamental Symmetries

• Lepton number & Sterile Neutrinos



 $0\nu\beta\beta$ LEGEND (ORNL lead US lab)

Connecting nuclei to the cosmos

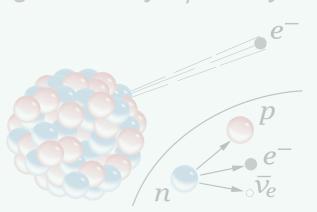
• Dense matter & Neutron stars





Precision frontier

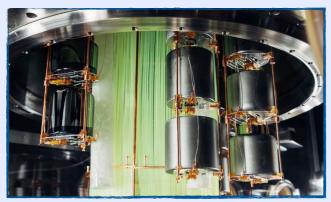
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Connecting nuclei to the cosmos

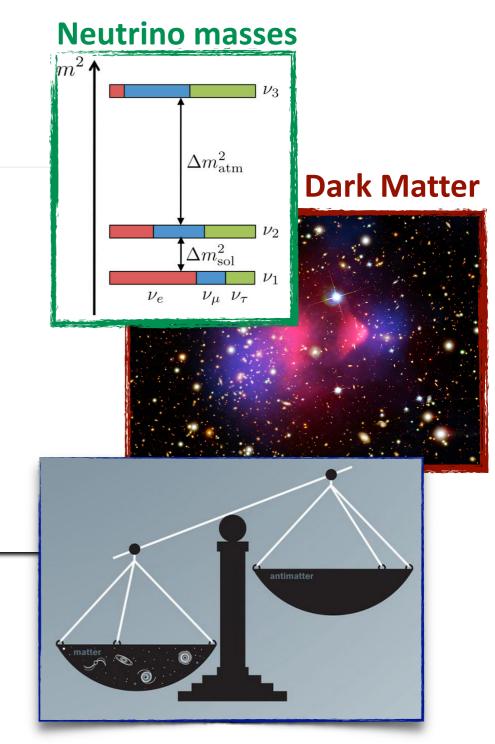
Dense matter & Neutron stars



Sterile neutrinos & $0\nu\beta\beta$

- Sterile neutrinos can answer open questions:
 - Neutrino masses
 - Dark matter candidate
 - Matter-antimatter Asymmetry

Appear in many BSM scenarios: Left-Right/GUTs...

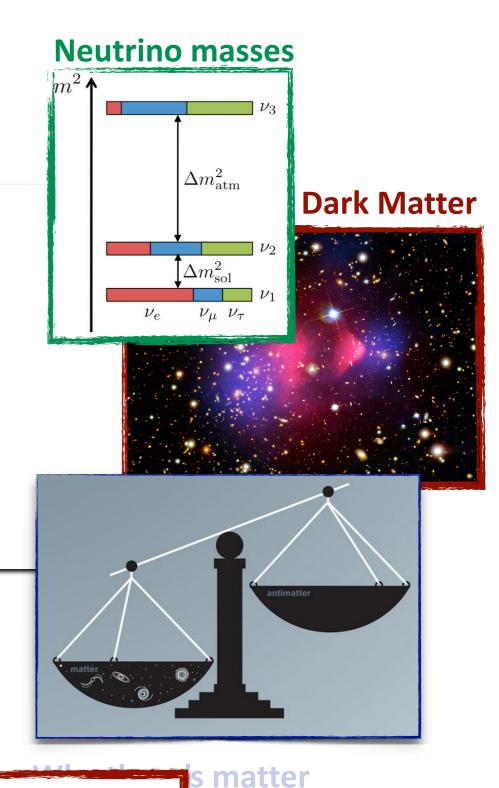


Why there's matter

Sterile neutrinos & $0\nu\beta\beta$

- Sterile neutrinos can answer open questions:
 - Neutrino masses
 - Dark matter candidate
 - Matter-antimatter Asymmetry

Appear in many BSM scenarios: Left-Right/GUTs...



Generally leads to Lepton-Number Violation

$$\Longrightarrow 0\nu\beta\beta$$

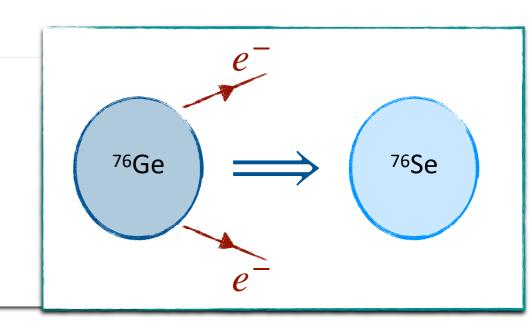
Neutrinoless Double beta decay

What & why?

What is $0\nu\beta\beta$?

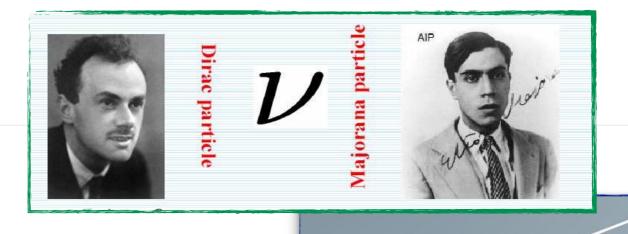
• $A(Z,N) \to A(Z+2,N-2) + 2e^- + 0\bar{\nu}$

Lepton Number Violating (LNV)



Why look for $0\nu\beta\beta$?

- Detection sheds light on:
 - Neutrino masses
 - ullet Observation would show that u's are Majorana particles
 - Matter-antimatter asymmetry



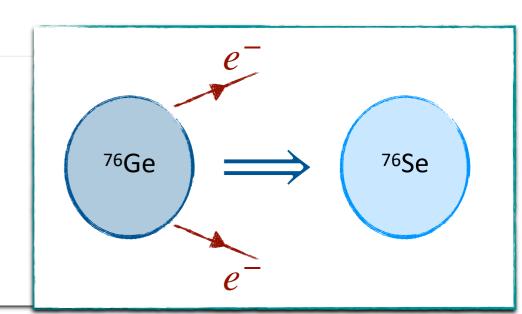
Neutrinoless Double beta decay

What & why?

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• $A(Z,N) \to A(Z+2,N-2) + 2e^- + 0\bar{\nu}$

Lepton Number Violating (LNV)

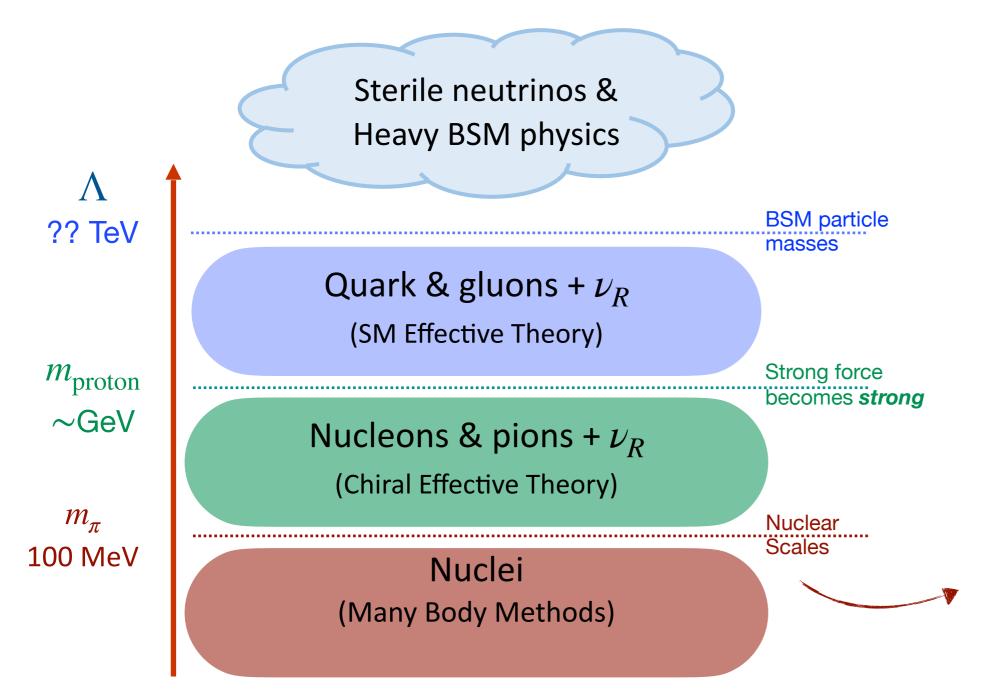


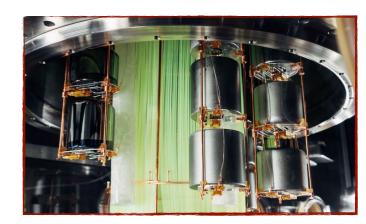
- Large scale experimental program underway
 - Highest priority in Long Range Plan for Nuclear Science
- Theory is needed to connect to the underlying physics
 - · Sterile neutrinos



Developed EFT framework for ν_R in $0\nu\beta\beta$

From BSM scales to nuclear scales





 $0\nu\beta\beta$ LEGEND (ORNL US lead lab)

Developed EFT framework for ν_R in $0\nu\beta\beta$

From BSM scales to nuclear scales

Sterile neutrinos & Heavy BSM physics

vy BSM physics

Quark & gluons + ν_R (SM Effective Theory)

Nucleons & pions + ν_R (Chiral Effective Theory)

Nuclei (Many Body Methods) PHYSICAL REVIEW C 108, 045501 (2023)

Neutrinoless double- β decay in the neutrino-extended standard model

Wouter Dekens,^{1,*} Jordy Pat

BSM particle

Strong force

Nuclear Scales

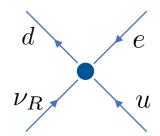
becomes strong

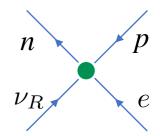
masses

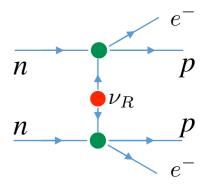
RECEIVED: March 7, 2020
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ACCEPTED: May 19, 2020
PUBLISHED: June 16, 2020

Sterile neutrinos and neutrinoless double beta decay in effective field theory

W. Dekens, a J. de Vries, b,c K. Fuyuto, b,d E. Mereghetti d and G. Zhou b







 $m_{
m proton}$

?? TeV

~GeV

 m_π 100 MeV

Neutrinoless double beta decay

Example for heavy ν_R : Majorana-mass mechanism

- **Fundamental Physics**
 - Effective neutrino mass

 $\Gamma_{0
uetaeta} \propto M_{
u}^{2} m_{etaeta}^{2}$

Theory input: Nuclear Matrix Element

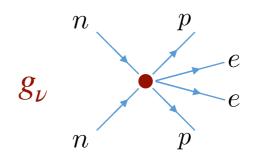
 $M_{
u} \sim \frac{p}{n}$ pPrevious

literature

PHYSICAL REVIEW LETTERS 120, 202001 (2018)

New Leading Contribution to Neutrinoless Double-β Decay

Vincenzo Cirigliano,¹ Wouter Dekens,¹ Jordy de Vries,² Michael L. Graesser,¹ Emanuele Mereghetti,¹ Saori Pastore,¹ and Ubirajara van Kolck^{3,4}



New interaction Needed at *leading order*

- Provided first estimate of $g_{
u}$

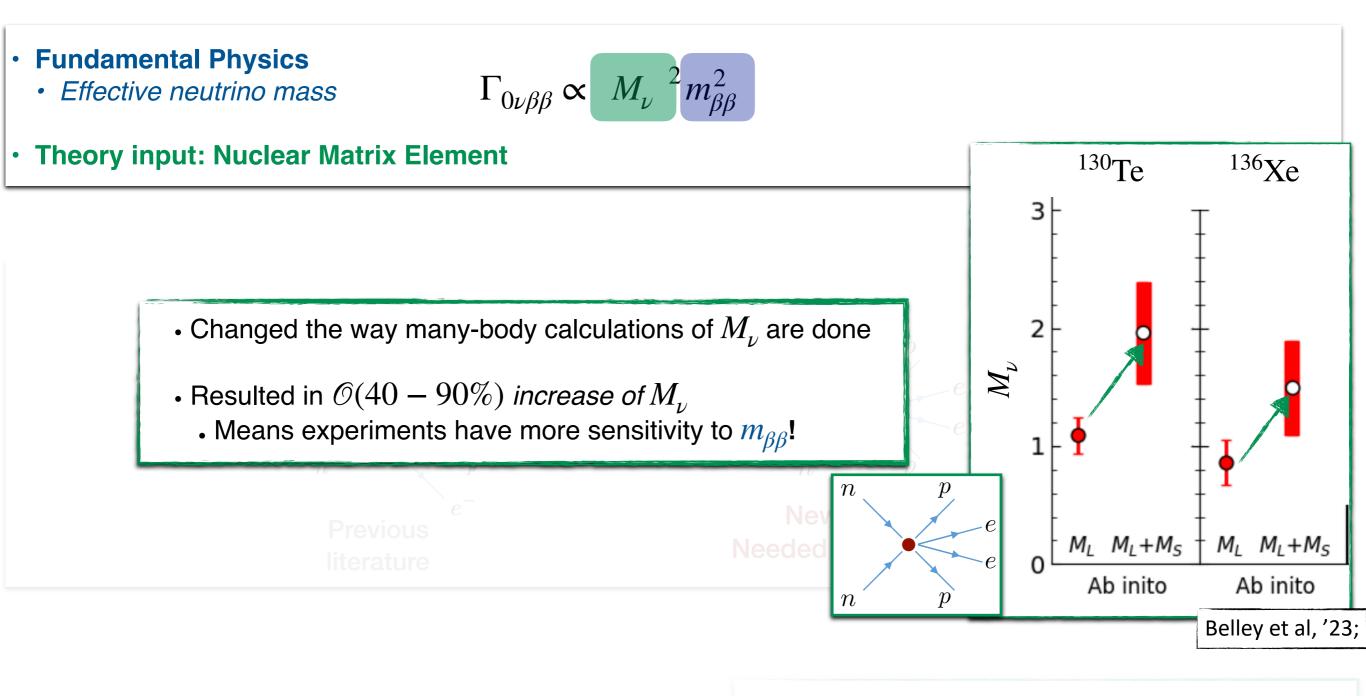
PHYSICAL REVIEW LETTERS 126, 172002 (2021)

Toward Complete Leading-Order Predictions for Neutrinoless Double β Decay

Vincenzo Cirigliano[®], Wouter Dekens[®], Jordy de Vries[®], Martin Hoferichter[®], and Emanuele Mereghetti[®]1

Neutrinoless double beta decay

Example for heavy ν_R : Majorana-mass mechanism



ullet Provided first estimate of $g_{
u}$

PHYSICAL REVIEW LETTERS 126, 172002 (202

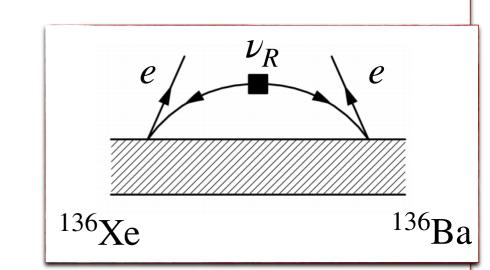
Toward Complete Leading-Order Predictions for Neutrinoless Double β Decay

/incenzo Cirigliano^{®, 1} Wouter Dekens^{®, 2} Jordy de Vries^{®, 3,4,5,6} Martin Hoferichter^{®, 7} and Emanuele Mereghetti

Neutrinoless double beta decay

Example: Light ν_R

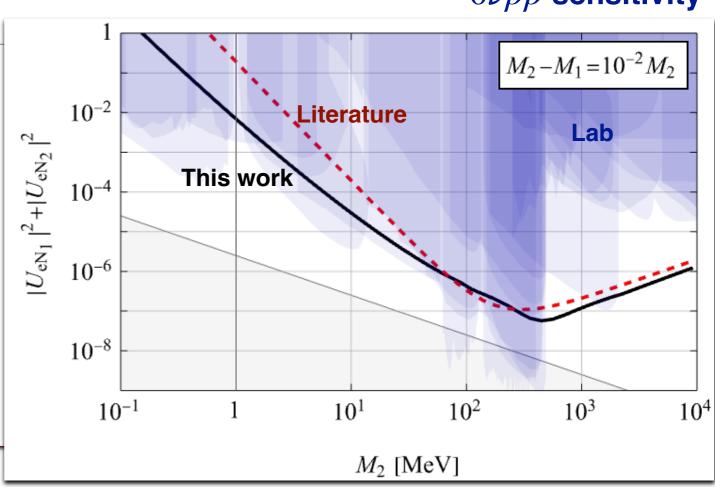
- Need to keep sterile neutrinos (ν_R) in the hadronic theory
- Leads to new effects from ν_R with small momenta $\,\sim \, MeV$
 - These ν_R couple to the nucleus as a whole
- This contribution was missed in the literature



$0 u\beta\beta$ sensitivity

- Impact in a toy model
 - "1+1+1" scenario

- Is the leading contribution for light ν_R
- Systematic EFT description crucial



Fundamental Symmetries Future projects

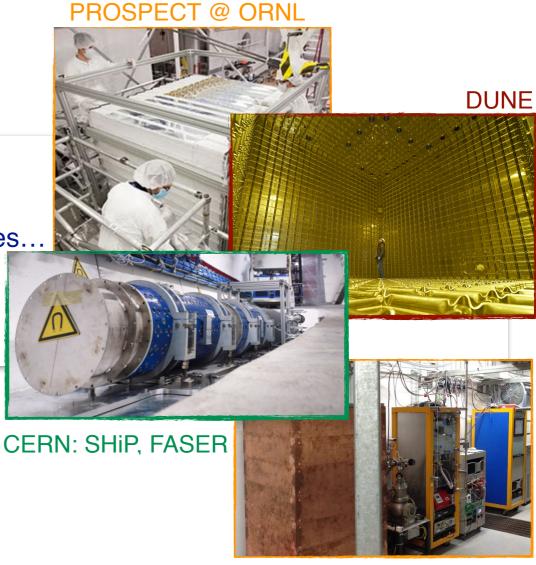
Lepton Number Violation and Sterile neutrinos Future work

Future work

• Combine $0\nu\beta\beta$ with other probes in the same EFT framework

• Colliders, reactor, oscillation experiments, astrophysical probes...

Needed to assess viability of sterile-neutrino scenarios



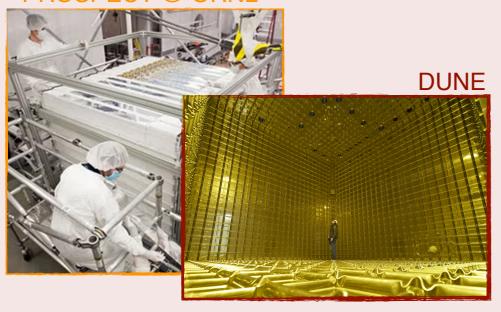
COHERENT @ ORNL

Lepton number & Sterile Neutrinos

Global Analysis of ν_R interactions

• Combine $0\nu\beta\beta$ with Colliders, reactor, oscillation experiments, astrophysical probes...

PROSPECT @ ORNL



Overlap with the Astrophysics & HEP Groups

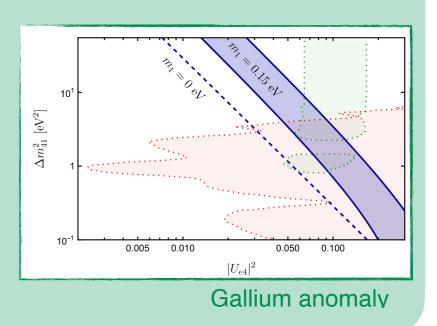
ν cross sections in reactor experiments

Currently involve anomalies

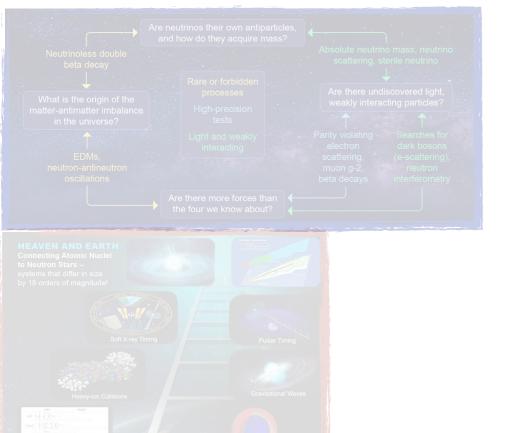
• Can be interpreted as light sterile neutrinos

SM prediction using Chiral EFT framework

• Interplay with "Precision Frontier" direction

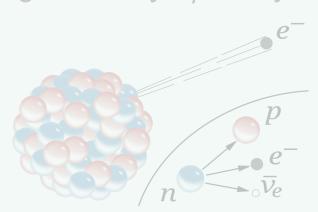


Overlap with the Nuclear Theory Group



Precision frontier

ullet Cabibbo Angle Anomaly & eta decays



Effectively probing (Beyond) the SM physics

Fundamental Symmetries

• Lepton number & Sterile Neutrinos

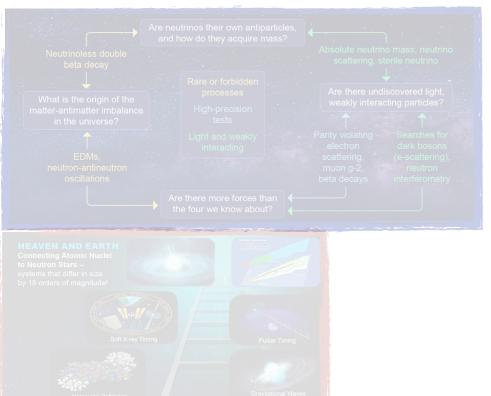


 $0\nu\beta\beta$ LEGEND (ORNL lead US lab)

Connecting nuclei to the cosmos

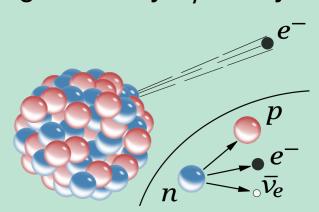
Dense matter & Neutron stars





Precision frontier

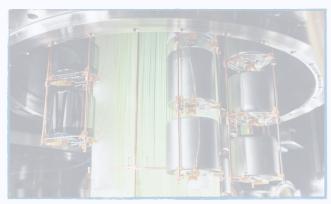
• Cabibbo Angle Anomaly & β decays



Effectively probing (Beyond) the SM physics

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 $0\nu\beta\beta$ LEGEND (ORNL lead US lab)

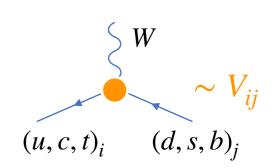
Connecting nuclei to the cosmos

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β decays

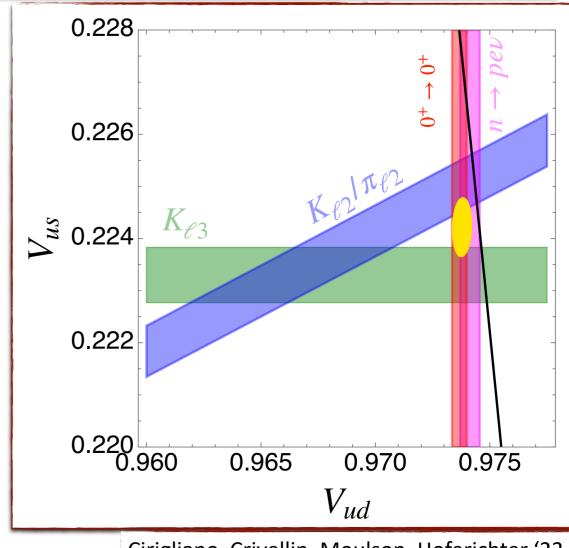
- Helped build the SM, probe ${\it O}(10)$ TeV BSM scales
- Determine V_{ud} very precisely $\delta V_{ud} \sim {\rm few} \times 10^{-4}$



• CKM unitarity test: $\Delta = V_{ud}^2 + V_{us}^2 + V_{ub}^2 - 1 = 0$

Unitarity test

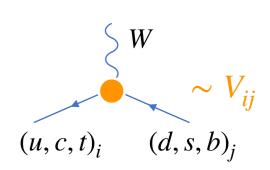
- Experimentally determined
 - V_{ud} from neutron & nuclear eta decays
 - ullet V_{us} from Kaon/pion decays
- Disagrees with unitarity by $\sim 3\sigma$
 - Issue with SM predictions/uncertainties?
 - BSM physics?



Cirigliano, Crivellin, Moulson, Hoferichter '23

β decays

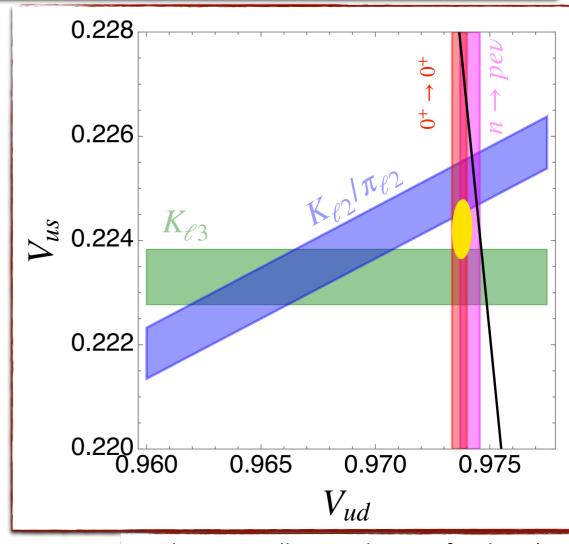
- Helped build the SM, probe ${\it O}(10)$ TeV BSM scales
- Determine V_{ud} very precisely $\delta V_{ud} \sim {\rm few} \times 10^{-4}$



• CKM unitarity test: $\Delta = V_{ud}^2 + V_{us}^2 + V_{ub}^2 - 1 = 0$

Unitarity test

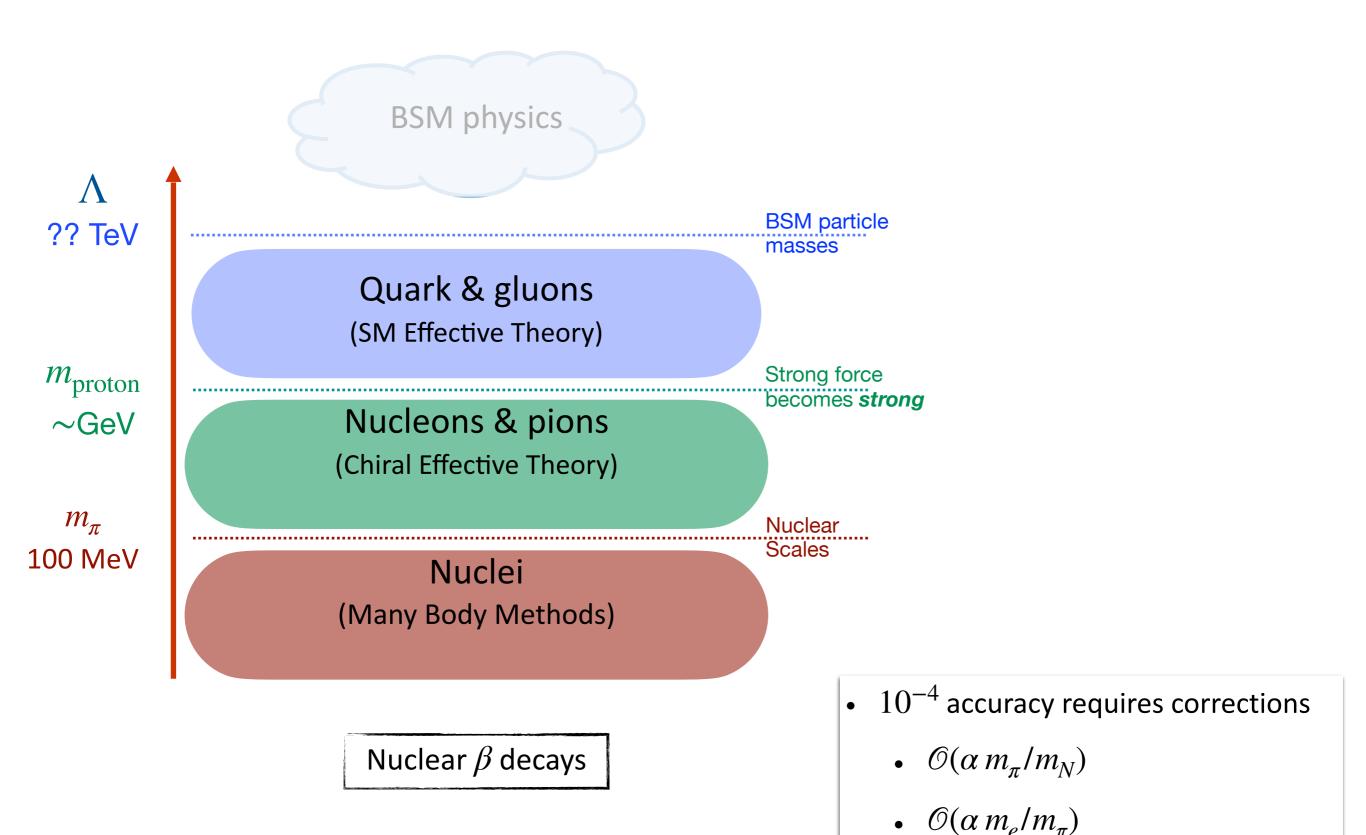
- Experimentally determined
 - V_{ud} from neutron & nuclear eta decays
 - ullet V_{us} from Kaon/pion decays
- Disagrees with unitarity by $\sim 3\sigma$
 - Issue with SM predictions/uncertainties?
 - BSM physics?



Cirigliano, Crivellin, Moulson, Hoferichter '23

Developed EFT framework for nuclear β decays

From BSM scales to nuclear scales

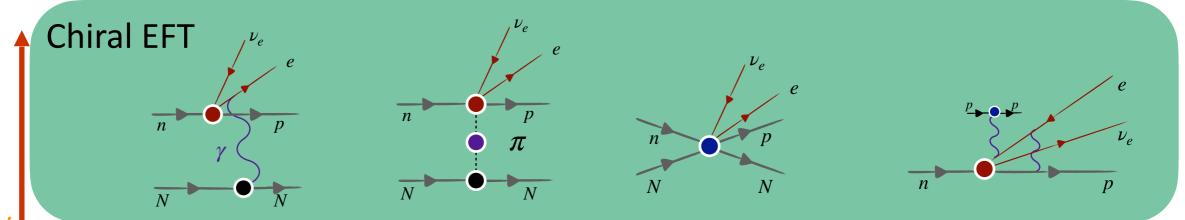


Developed EFT framework for one & two nucleon effects

PHYSICAL REVIEW LETTERS 133, 211801 (2024)

Radiative Corrections to Superallowed β Decays in Effective Field Theory

Vincenzo Cirigliano[®], Wouter Dekens[®], Jordy de Vries[®], Stefano Gandolfi[®], 4



100 MeV

 $\mathcal{O}(\alpha m_\pi/m_N)$ & $\mathcal{O}(\alpha m_e/m_\pi)$ two-nucleon terms

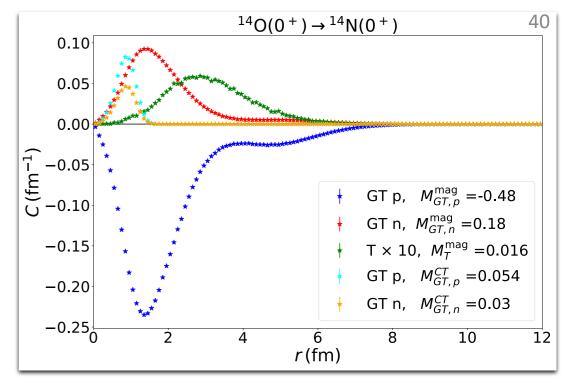
 $\mathcal{O}(\alpha^2 Z)$ two-nucleon terms

Nuclear-structure dependence, " $\delta_{\!N\!S}$ "

- Source of largest uncertainty
- In EFT described by $\langle \psi_i \ V \ \psi_i \rangle$

Half Life

 $^{14}O \rightarrow ^{14}N$



Nuclear Matrix Elements

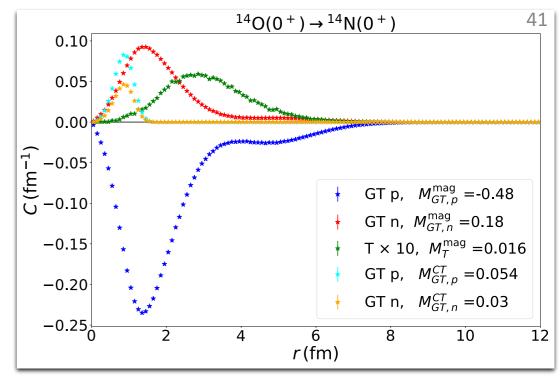
• All combined:
$$V_{ud} = 0.97364(12)_{g_V}(10)_{\exp}(22)_{\bar{f}}(13)_{\delta_{NS}^{\text{non-LEC}}}(44)_{\delta_{NS}^{\text{LEC}}}(12)_{\delta_c}[55]_{\text{total}}$$

ullet Compatible with traditional approach $V_{ud} = 0.97405(37)_{\mathrm{total}}$

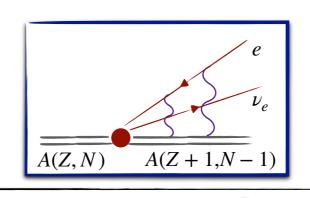
Hardy & Towner, '20

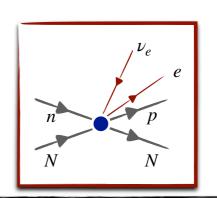
Half Life

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Nuclear Matrix Elements





- All combined: $V_{ud} = 0.97364(12)_{g_V}(10)_{\text{exp}}(22)_{\overline{f}}(13)_{\delta_{NS}^{\text{non-LEC}}}(44)_{\delta_{NS}^{\text{LEC}}}(12)_{\delta_c}[55]_{\text{total}}$
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Hardy & Towner, '20

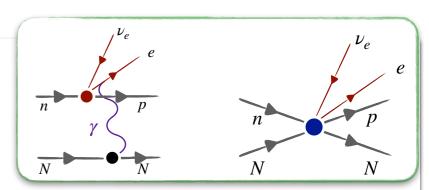
But larger uncertainty

Precision frontier Future projects

Precision frontier Future work

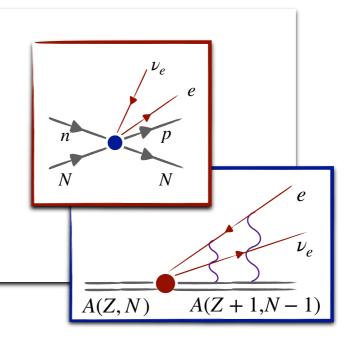
Controlling theory uncertainties

- Nuclear β decays measured to $\mathcal{O}(10^{-4})$ precision
- Obtain same level of precision using EFT framework
 - Needed to unlock the full BSM reach of experiments



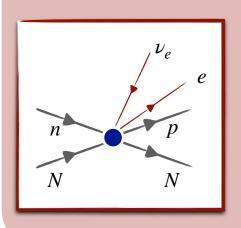
Radiative corrections in EFT

- Requires control of:
 - Unknown short-distance interactions
 - Missing $O(\alpha^2 Z)$ contributions
- EFT allows for systematic improvements



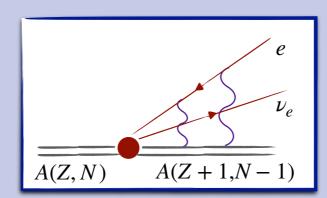
Short-distance couplings

- Fit to measurements
- Obtain from theory



$O(\alpha^2 Z)$ corrections

• Compute two-loop diagrams



Isospin breaking effects

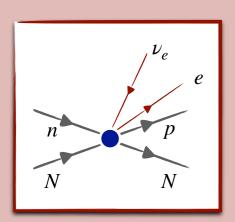
Develop EFT framework for " δ_{C} "

$$\langle \psi_f \ \tau^+ \ \psi_i \rangle = \sqrt{2}(1 - \delta_C/2)$$

Overlap with Nuclear Theory Group

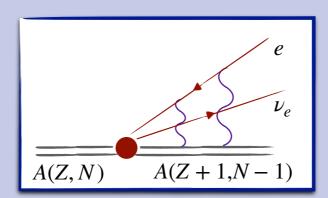
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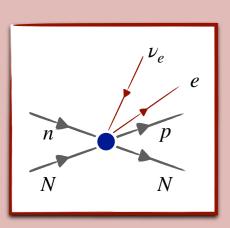
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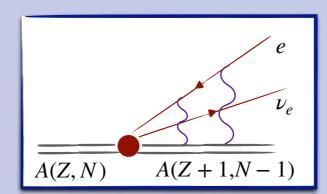
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Isospin breaking effects

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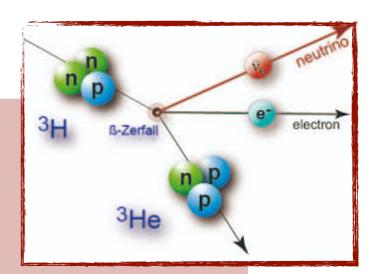
$$\langle \psi_f \ \tau^+ \ \psi_i \rangle = \sqrt{2}(1 - \delta_C/2)$$

NTNP topical collaboration

Spin off: beyond $0^+ \rightarrow 0^+$ decays

Generalize to decays of tritium

- Help determine three-nucleon forces
 - Could be affected by radiative corrections
 - Interplay with "Connecting nuclei to the Cosmos"



Funding strategies

Funding strategies

Initially

- Possible start-up funds
 - A student & postdoc
- Apply for
 - DOE: Early Career Award, regular FOAs
 - NSF: CAREER program, regular FOAs



Other sources of funding

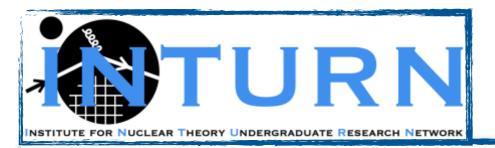
- Lead and/or participate in collaborative grant applications
 - DOE topical collaborations
 - NSF hubs
- Explore
 - Departmental grants
 - RENEW / FAIR
- Sources of funding in related fields
- E.g. machine learning has applications to
 - BSM fits for sterile neutrinos
 - Effects of three-nucleon forces in Neutron Stars

Mentoring

- Recent/current mentorship
 - Thomas Cepollina (undergraduate)
 - Maria Dawid (Graduate student)
 - Sebastián Urrutia Quiroga (Postdoc)
 - Heleen Mulder (Master student, Nikhef/Amsterdam)



Maria Dawid



- The INT's Undergraduate Research Network
 - Provides research opportunities to undergraduates from diverse backgrounds
 - Aimed at ensuring they are able to stay in physics
- Part of organization/selection committee
- I currently mentor Thomas Cepollina (3rd-year student) on a research project

At UTK

- Act as a mentor in the Nuclear Physics Eastern Tennessee Fellowship
- Outreach: High School Lecture Series

- Exciting portfolio of (upcoming) experiments
- Will probe
 - Beyond the SM physics to
 - Nuclear physics of dense matter
- Nuclear theory is needed to
 - Connect different measurements
 - Unlock their full BSM reach

- Addressed by research program with 3 pillars
 - Sterile neutrinos & Lepton number
 - $\cdot \beta$ decays & the Cabibbo Angle
 - **Connecting Nuclei to Neutron Stars**

Each has opportunities for collaboration with

- UTK Nuclear Theory
- UTK Astrophysics & HEP
- NTNP

